



**OPTICAL COEFFICIENTS IN THE N(P)-TYPE DEGENERATE  
GaTe(1-x) P(x)-CRYSTALLINE ALLOY, DUE TO THE NEW STATIC  
DIELECTRIC CONSTANT-LAW AND THE GENERALIZED MOTT  
CRITERIUM IN THE METAL-INSULATOR TRANSITION (8)**

**Prof. Dr. Huynh Van Cong\***

Université de Perpignan Via Domitia, Laboratoire de Mathématiques et Physique (LAMPS),  
EA 4217, Département de Physique, 52, Avenue Paul Alduy, F-66 860 Perpignan, France.

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**\*Corresponding Author**

**Prof. Dr. Huynh Van  
Cong**

Université de Perpignan Via  
Domitia, Laboratoire de  
Mathématiques et Physique  
(LAMPS), EA 4217,  
Département de Physique,  
52, Avenue Paul Alduy, F-  
66 860 Perpignan, France.

**ABSTRACT**

In the n(p)-type  $\text{GaTe}_{1-x}\text{P}_x$ - crystalline alloy, with  $0 \leq x \leq 1$ , basing on our two recent works<sup>[1, 2]</sup>, for a given  $x$ , and with an increasing  $r_{d(a)}$ , the optical coefficients have been determined, as functions of the photon energy  $E$ , total impurity density  $N$ , the donor (acceptor) radius  $r_{d(a)}$ , concentration  $x$ , and temperature  $T$ . Those results have been affected by (i) the important new  $\varepsilon(r_{d(a)}, x)$ -law, developed in Equations (8a, 8b), stating that, for a given  $x$ , due to the impurity-size effect,  $\varepsilon$  decreases ( $\searrow$ ) with an increasing ( $\nearrow$ )  $r_{d(a)}$ , and then by (ii) the generalized Mott critical  $d(a)$ -density defined in the metal-insulator transition (MIT),  $N_{CDn(NDp)}(r_{d(a)}, x)$ , as observed in Equations (8c, 9a). Furthermore, we also showed that  $N_{CDn(NDp)}$  is just the density of

carriers localized in exponential band tails, with a precision of the order of  $2.91 \times 10^{-7}$ , as that given in Table 4 of Ref.<sup>[1]</sup>, according to a definition of the effective density of electrons (holes) given in parabolic conduction (valence) bands by:  $N^*(N, r_{d(a)}, x) \equiv N - N_{CDn(NDp)}(r_{d(a)}, x)$ , as defined in Eq. (9d). In summary, due to the new  $\varepsilon(r_{d(a)}, x)$ -law and to the effective density of electrons (holes) given in parabolic conduction (valence) bands  $N^*(N, r_{d(a)}, x)$ , for a given  $x$ , and with an increasing  $r_{d(a)}$ , the numerical results of all the optical coefficients, obtained in appropriated physical conditions ( $E, N, T$ ),

and calculated by using Equations (15, 16, 20, 21), are reported in Tables 1, 2, 3n, 3p, 4n, 4p, 5n, and 5p in Appendix 1.

**KEYWORDS:**  $\text{GaTe}_{1-x}\text{P}_x$ - crystalline alloy; impurity-size effect; Mott critical impurity density in the MIT, optical coefficients.

## INTRODUCTION

Here, basing on our two recent works<sup>[1, 2]</sup> and also other ones<sup>[3-8]</sup>, all the optical coefficients given in the n(p)-type  $\mathbf{X(x)} \equiv \text{GaTe}_{1-x}\text{P}_x$  - crystalline alloy, with  $0 \leq x \leq 1$ , are investigated, as functions of the photon energy E, total impurity density N, the donor (acceptor) radius  $r_{d(a)}$ , concentration x, and temperature T.

Then, for a given x, and with an increasing  $r_{d(a)}$ , the numerical results of all the optical coefficients, obtained in appropriated physical conditions (E, N, T), and calculated by using Equations (15, 16, 20, 21), are reported in Tables 1, 2, 3n, 3p, 4n, 4p, 5n, and 5p in Appendix 1.

## ENERGY BAND STRUCTURE PARAMETERS

First of all, in the  $n^+(p^+) - p(n)$   $\mathbf{X(x)}$ - crystalline alloy at  $T=0$  K, we denote the donor (acceptor) d(a)-radius by  $r_{d(a)}$ , and also the intrinsic one by:  $r_{do(ao)} = r_{Te(Ga)} = 0.132$  nm (0.126 nm).

### A. Effect of x- concentration

Here, the intrinsic energy-band-structure parameters<sup>[1]</sup>, are expressed as functions of x, are given in the following.

(i)-The unperturbed relative effective electron (hole) mass in conduction (valence) bands are given by:

$$m_{c(v)}(x)/m_o = 0.13 (0.5) \times x + 0.209 (0.4) \times (1 - x) \quad (1)$$

(ii)-The unperturbed relative static dielectric constant of the intrinsic of the single crystalline X- alloy is found to be defined by:

$$\epsilon_o(x) = 11.1 \times x + 12.3 \times (1 - x). \quad (2)$$

(iii)-Finally, the unperturbed band gap at 0 K is found to be given by:

$$E_{go}(x) = 1.796 \times x + 1.796 \times (1 - x) = 1.796. \quad (3)$$

Therefore, we can define the effective donor (acceptor)-ionization energy in absolute values as:

$$E_{\text{do(ao)}}(x) = \frac{13600 \times [m_{\text{c(v)}}(x)/m_0]}{[\varepsilon_0(x)]^2} \text{ meV}, \quad (4)$$

and then, the isothermal bulk modulus, by:

$$B_{\text{do(ao)}}(x) \equiv \frac{E_{\text{do(ao)}}(x)}{\left(\frac{4\pi}{3}\right) \times (r_{\text{do(ao)}})^3}. \quad (5)$$

### B. Effect of Impurity $r_{\text{d(a)}}$ -size, with a given $x$

Here, the changes in all the energy-band-structure parameters, expressed in terms of the effective relative dielectric constant  $\varepsilon(r_{\text{d(a)}}, x)$ , developed as follows.

At  $r_{\text{d(a)}} = r_{\text{do(ao)}}$ , the needed boundary conditions are found to be, for the impurity-atom volume  $V = (4\pi/3) \times (r_{\text{d(a)}})^3$ ,  $V_{\text{do(ao)}} = (4\pi/3) \times (r_{\text{do(ao)}})^3$ , for the pressure  $p$ ,  $p_0 = 0$ , and for the deformation potential energy (or the strain energy)  $\sigma$ ,  $\sigma_0 = 0$ . Further, the two important equations<sup>[1, 7]</sup>, used to determine the  $\sigma$ -variation,  $\Delta\sigma \equiv \sigma - \sigma_0 = \sigma$ , are defined by:

$\frac{dp}{dV} = -\frac{B}{V}$  and  $p = -\frac{d\sigma}{dV}$ . giving:  $\frac{d}{dV}\left(\frac{d\sigma}{dV}\right) = \frac{B}{V}$ . Then, by an integration, one gets:

$$\begin{aligned} [\Delta\sigma(r_{\text{d(a)}}, x)]_{\text{n(p)}} &= B_{\text{do(ao)}}(x) \times (V - V_{\text{do(ao)}}) \times \ln \\ \left(\frac{V}{V_{\text{do(ao)}}}\right) &= E_{\text{do(ao)}}(x) \times \left[ \left(\frac{r_{\text{d(a)}}}{r_{\text{do(ao)}}}\right)^3 - 1 \right] \times \ln \left(\frac{r_{\text{d(a)}}}{r_{\text{do(ao)}}}\right)^3 \geq 0. \end{aligned} \quad (6)$$

Furthermore, we also shown that, as  $r_{\text{d(a)}} > r_{\text{do(ao)}}$  ( $r_{\text{d(a)}} < r_{\text{do(ao)}}$ ), the compression (dilatation) gives rise to the increase (the decrease) in the energy gap  $E_{\text{gn(gp)}}(r_{\text{d(a)}}, x)$ , and the effective donor (acceptor)-ionization energy  $E_{\text{d(a)}}(r_{\text{d(a)}}, x)$  in absolute values, obtained in the effective Bohr model, which is represented respectively by:  $\pm [\Delta\sigma(r_{\text{d(a)}}, x)]_{\text{n(p)}}$ ,

$$E_{\text{gno(gp)}}(r_{\text{d(a)}}, x) - E_{\text{go}}(x) = E_{\text{d(a)}}(r_{\text{d(a)}}, x) - E_{\text{do(ao)}}(x) = E_{\text{do(ao)}}(x) \times \left[ \left(\frac{\varepsilon_0(x)}{\varepsilon(r_{\text{d(a)}})}\right)^2 - 1 \right] = + [\Delta\sigma(r_{\text{d(a)}}, x)]_{\text{n(p)}}$$

for  $r_{\text{d(a)}} \geq r_{\text{do(ao)}}$ , and for  $r_{\text{d(a)}} \leq r_{\text{do(ao)}}$ ,

$$E_{\text{gno(gp)}}(r_{\text{d(a)}}, x) - E_{\text{go}}(x) = E_{\text{d(a)}}(r_{\text{d(a)}}, x) - E_{\text{do(ao)}}(x) = E_{\text{do(ao)}}(x) \times \left[ \left(\frac{\varepsilon_0(x)}{\varepsilon(r_{\text{d(a)}})}\right)^2 - 1 \right] = - [\Delta\sigma(r_{\text{d(a)}}, x)]_{\text{n(p)}} \quad (7)$$

Therefore, from Equations (6) and (7), one obtains the expressions for relative dielectric constant  $\epsilon(r_{d(a)}, x)$  and energy band gap  $E_{gn(gp)}(r_{d(a)}, x)$ , as:

(i)-for  $r_{d(a)} \geq r_{do(ao)}$ , since  $\epsilon(r_{d(a)}, x) = \frac{\epsilon_o(x)}{\sqrt{1 + \left[ \left( \frac{r_{d(a)}}{r_{do(ao)}} \right)^3 - 1 \right] \times \ln \left( \frac{r_{d(a)}}{r_{do(ao)}} \right)^3}} \leq \epsilon_o(x)$ , being a **new**

$\epsilon(r_{d(a)}, x)$ -law,

$$E_{gno(gp_o)}(r_{d(a)}, x) - E_{go}(x) = E_{d(a)}(r_{d(a)}, x) - E_{do(ao)}(x) = E_{do(ao)}(x) \times \left[ \left( \frac{r_{d(a)}}{r_{do(ao)}} \right)^3 - 1 \right] \times \ln \left( \frac{r_{d(a)}}{r_{do(ao)}} \right)^3 \geq 0, \tag{8a}$$

according to the increase in both  $E_{gn(gp)}(r_{d(a)}, x)$  and  $E_{d(a)}(r_{d(a)}, x)$ , with increasing  $r_{d(a)}$  and for a given  $x$ , and

(ii)-for  $r_{d(a)} \leq r_{do(ao)}$ , since  $\epsilon(r_{d(a)}, x) = \frac{\epsilon_o(x)}{\sqrt{1 - \left[ \left( \frac{r_{d(a)}}{r_{do(ao)}} \right)^3 - 1 \right] \times \ln \left( \frac{r_{d(a)}}{r_{do(ao)}} \right)^3}} \geq \epsilon_o(x)$ , with a

condition, given by:  $\left[ \left( \frac{r_{d(a)}}{r_{do(ao)}} \right)^3 - 1 \right] \times \ln \left( \frac{r_{d(a)}}{r_{do(ao)}} \right)^3 < 1$ , being a **new**  $\epsilon(r_{d(a)}, x)$ -law,

$$E_{gno(gp_o)}(r_{d(a)}, x) - E_{go}(x) = E_{d(a)}(r_{d(a)}, x) - E_{do(ao)}(x) = -E_{do(ao)}(x) \times \left[ \left( \frac{r_{d(a)}}{r_{do(ao)}} \right)^3 - 1 \right] \times \ln \left( \frac{r_{d(a)}}{r_{do(ao)}} \right)^3 \leq 0, \tag{8b}$$

corresponding to the decrease in both  $E_{gn(gp)}(r_{d(a)}, x)$  and  $E_{d(a)}(r_{d(a)}, x)$ , with decreasing  $r_{d(a)}$  and for a given  $x$ ; therefore, the effective Bohr radius  $a_{Bn(Bp)}(r_{d(a)}, x)$  is defined by:

$$a_{Bn(Bp)}(r_{d(a)}, x) \equiv \frac{\epsilon(r_{d(a)}, x) \times \hbar^2}{m_{c(v)}(x) \times q^2} = 0.53 \times 10^{-8} \text{ cm} \times \frac{\epsilon(r_{d(a)}, x)}{m_{c(v)}(x)/m_o}. \tag{8c}$$

Furthermore, it is interesting to remark that the critical total donor (acceptor)-density in the metal-insulator transition (MIT) at  $T=0$  K,  $N_{CDn(NDp)}(r_{d(a)}, x)$ , was given by the Mott's criterium, with an empirical parameter,  $M_{n(p)}$ , as:

$$N_{CDn(NDp)}(r_{d(a)}, x)^{1/3} \times a_{Bn(Bp)}(r_{d(a)}, x) = M_{n(p)}, \quad M_{n(p)} = 0.25, \tag{9a}$$

depending thus on our **new**  $\epsilon(r_{d(a)}, x)$ -law.

This excellent one can be explained from the definition of the reduced effective Wigner-Seitz (WS) radius  $r_{sn(sp)}$ , characteristic of interactions, by:

$$r_{sn(sp)}(N, r_{d(a)}, x) \equiv \left( \frac{3}{4\pi N} \right)^{1/3} \times \frac{1}{a_{Bn(Bp)}(r_{d(a)}, x)} = 1.1723 \times 10^8 \times \left( \frac{1}{N} \right)^{1/3} \times \frac{m_{c(v)}(x)/m_o}{\epsilon(r_{d(a)}, x)}, \tag{9b}$$

being equal to, in particular, at  $N=N_{\text{CDn(CDP)}}(r_{\text{d(a)},x})$ :  $r_{\text{sn(sp)}}(N_{\text{CDn(CDP)}}(r_{\text{d(a)},x}), r_{\text{d(a)},x})=$   
**2.4814**, for any  $(r_{\text{d(a)},x})$ -values. So, from Eq. (9b), one also has:

$$N_{\text{CDn(CDP)}}(r_{\text{d(a)},x})^{1/3} \times a_{\text{Bn(Bp)}}(r_{\text{d(a)},x}) = \left(\frac{3}{4\pi}\right)^{1/3} \times \frac{1}{2.4814} = \mathbf{0.25} = (\mathbf{WS})_{\text{n(p)}} = \mathbf{M}_{\text{n(p)}}. \quad (9c)$$

Thus, the above Equations (9a, 9b, 9c) confirm our new  $\varepsilon(r_{\text{d(a)},x})$ -law, given in Equations (8a, 8b).

Furthermore, by using  $\mathbf{M}_{\text{n(p)}} = \mathbf{0.25}$ , according to the empirical Heisenberg parameter  $\mathcal{H}_{\text{n(p)}} = \mathbf{0.47137}$ , as those given in Equations (8, 15) of the Ref.<sup>[1]</sup>, we have also showed that  $N_{\text{CDn(CDP)}}$  is just the density of electrons (holes) localized in the exponential conduction (valence)-band tail, with a precision of the order of  $2.91 \times 10^{-7}$ . Therefore, the density of electrons (holes) given in parabolic conduction (valence) bands can be defined, as that given in compensated materials, by:

$$N^*(N, r_{\text{d(a)},x}) \equiv N - N_{\text{CDn(NDp)}}(r_{\text{d(a)},x}). \quad (9d)$$

### C. Effect of temperature T, with given x and $r_{\text{d(a)}}$

Here, the intrinsic band gap  $E_{\text{gni(gp)}}(r_{\text{d(a)},x}, T)$  at any T is given by:

$$E_{\text{gni(gp)}}(r_{\text{d(a)},x}, T) \text{ in eV} = E_{\text{gno(gp)}}(r_{\text{d(a)},x}) - 10^{-4} \times T^2 \times \left\{ \frac{5.405 \times x}{T+204 \text{ K}} + \frac{3.065 \times (1-x)}{T+94 \text{ K}} \right\}, \quad (10)$$

suggesting that, for given x and  $r_{\text{d(a)}}$ ,  $E_{\text{gni(gp)}}$  decreases with an increasing T.

Then, in the following, for the study of optical phenomena, one denote the conduction (valence)-band density of states by  $N_{\text{c(v)}}(T, x)$  as:

$$N_{\text{c(v)}}(T, x) = 2 \times g_{\text{c(v)}}(x) \times \left( \frac{m_{\text{r(x)}} \times k_{\text{B}} T}{2\pi\hbar^2} \right)^{3/2} \text{ (cm}^{-3}\text{)}, \quad g_{\text{v}}(x) \equiv 1 \times x + 1 \times (1-x) = 1, \quad (11)$$

where  $m_{\text{r}}(x)/m_0$  is the reduced effective mass  $m_{\text{r}}(x)/m_0$ , defined by :

$$m_{\text{r}}(x) \equiv [m_{\text{c}}(x) \times m_{\text{v}}(x)] / [m_{\text{c}}(x) + m_{\text{v}}(x)].$$

### D. Heavy Doping Effect, with given T, x and $r_{\text{d(a)}}$

Here, as given in our previous works<sup>[1, 2]</sup>, the Fermi energy  $E_{\text{Fn}}(-E_{\text{Fp}})$ , and the band gap narrowing are reported in the following.

First, the reduced Fermi energy  $\eta_{n(p)}$  or the Fermi energy  $E_{Fn}(-E_{Fp})$ , obtained for any T and any effective d(a)-density,  $N^*(N, r_{d(a)}, x) = N^*$ , defined in Eq. (9d), for a simplicity of presentation, being investigated in our previous paper<sup>[8]</sup>, with a precision of the order of  $2.11 \times 10^{-4}$ , is found to be given by:

$$\eta_{n(p)}(u) \equiv \frac{E_{Fn}(u)}{k_B T} \left( \frac{-E_{Fp}(u)}{k_B T} \right) = \frac{G(u) + Au^B F(u)}{1 + Au^B}, \quad A = 0.0005372 \text{ and } B = 4.82842262, \quad (12)$$

where u is the reduced electron density,  $u(N, r_{d(a)}, x, T) \equiv \frac{N^*}{N_{c(v)}(T, x)}$ ,

$$F(u) = au^{\frac{2}{3}} \left( 1 + bu^{-\frac{4}{3}} + cu^{-\frac{8}{3}} \right)^{-\frac{2}{3}}, \quad a = [(3\sqrt{\pi}/4) \times u]^{2/3}, \quad b = \frac{1}{8} \left( \frac{\pi}{a} \right)^2, \quad c = \frac{62.3739855}{1920} \left( \frac{\pi}{a} \right)^4,$$

and  $G(u) \simeq \ln(u) + 2^{-\frac{3}{2}} \times u \times e^{-du}$ ;  $d = 2^{3/2} \left[ \frac{1}{\sqrt{27}} - \frac{3}{16} \right] > 0$ . Therefore, from Eq. (12),

the Fermi energies are expressed as functions of variables : N,  $r_{d(a)}$ , x, and T.

Here, one notes that: (i) as  $u \gg 1$ , according to the HD [d(a)- X(x)- alloy] ER-case, or to the degenerate case, Eq. (12) is reduced to the function F(u), and in particular at T=0 and as  $N^* = 0$ , according to the metal-insulator transition (MIT), one has:  $+E_{Fn}(-E_{Fp}) = \frac{\hbar^2}{2 \times m_r(x)} \times (3\pi^2 N^*)^{2/3} = 0$ , and (ii)  $\frac{E_{Fn}(u \ll 1)}{k_B T} \left( \frac{-E_{Fp}(u \ll 1)}{k_B T} \right) \ll -1$ , to the LD [a(d)- X(x)- alloy] BR-case, or to the non-degenerate case, Eq. (12) is reduced to the function G(u), noting that the notations: **HD(LD)** and **ER(BR)** denote the heavily doped (lightly doped)-cases and emitter (base)-regions, respectively.

Now, in Eq. (9b), in which one replaces  $m_{c(v)}(x)$  by  $m_r(x)$ , the effective Wigner-Seitz radius becomes as:

$$r_{sn(sp)}(N, r_{d(a)}, x) = 1.1723 \times 10^8 \times \left( \frac{g_{c(v)}(x)}{N^*} \right)^{1/3} \times \frac{m_r(x)}{\varepsilon(r_{d(a)}, x)}, \quad (13a)$$

the correlation energy of an effective electron gas,  $E_{cn(cp)}(N, r_{d(a)}, x)$ , is given as:

$$E_{cn(cp)}(N, r_{d(a)}, x) = \frac{-0.87553}{0.0908 + r_{sn(sp)}} + \frac{\frac{0.87553}{0.0908 + r_{sn(sp)}} + \left( \frac{2[1 - \ln(2)]}{\pi^2} \right) \times \ln(r_{sn(sp)}) - 0.093288}{1 + 0.03847728 \times r_{sn(sp)}^{1.67378876}}. \quad (13b)$$

Then, taking into account various spin-polarized chemical potential-energy contributions such as: exchange energy of an effective electron (hole) gas, majority-carrier correlation energy of an effective electron (hole) gas, minority hole (electron) correlation energy, majority electron (hole)-ionized d(a) interaction screened Coulomb potential energy, and



finally minority hole (electron)-ionized d(a) interaction screened Coulomb potential energy, the band gap narrowings are given in the following.

In the n-type HD X(x)- alloy, the BGN is found to be given by:

$$\Delta E_{\text{gno}}(N, r_d, x) = a_1 \times \frac{\varepsilon_0(x)}{\varepsilon(r_d, x)} \times N_r^{1/3} + a_2 \times \frac{\varepsilon_0(x)}{\varepsilon(r_d, x)} \times N_r^{1/3} \times (2.503 \times [-E_{\text{cn}}(r_{\text{sn}}) \times r_{\text{sn}}]) + a_3 \times \left[ \frac{\varepsilon_0(x)}{\varepsilon(r_d, x)} \right]^{5/4} \times \sqrt{\frac{m_v}{m_r}} \times N_r^{1/4} + a_4 \times \sqrt{\frac{\varepsilon_0(x)}{\varepsilon(r_d, x)}} \times N_r^{1/2} \times 2 + a_5 \times \left[ \frac{\varepsilon_0(x)}{\varepsilon(r_d, x)} \right]^2 \times N_r^{1/6}$$

$$, N_r \equiv \left( \frac{N^*}{N_{\text{CDn}}(r_d, x)} \right),$$

$$\Delta E_{\text{gn}}(N, r_d, x) = \Delta E_{\text{gno}}(N, r_d, x) \times \{2 \times x + 6 \times (1 - x)\}, \tag{14n}$$

where  $a_1 = 3.8 \times 10^{-3}(\text{eV})$  ,  $a_2 = 6.5 \times 10^{-4}(\text{eV})$  ,  $a_3 = 2.8 \times 10^{-3}(\text{eV})$  ,  $a_4 = 5.597 \times 10^{-3}(\text{eV})$  and  $a_5 = 8.1 \times 10^{-4}(\text{eV})$ , and in the p-type HD X(x)- alloy, as:

$$\Delta E_{\text{gpo}}(N, r_a, x) = a_1 \times \frac{\varepsilon_0(x)}{\varepsilon(r_a, x)} \times N_r^{1/3} + a_2 \times \frac{\varepsilon_0(x)}{\varepsilon(r_a, x)} \times N_r^{1/3} \times (2.503 \times [-E_{\text{cp}}(r_{\text{sp}}) \times r_{\text{sp}}]) + a_3 \times \left[ \frac{\varepsilon_0(x)}{\varepsilon(r_a, x)} \right]^{5/4} \times \sqrt{\frac{m_c}{m_r}} \times N_r^{1/4} + 2a_4 \times \sqrt{\frac{\varepsilon_0(x)}{\varepsilon(r_a, x)}} \times N_r^{1/2} + a_5 \times \left[ \frac{\varepsilon_0(x)}{\varepsilon(r_a, x)} \right]^2 \times N_r^{1/6}$$

$$, N_r \equiv \left( \frac{N^*}{N_{\text{CDp}}(r_a, x)} \right),$$

$$\Delta E_{\text{gp}}(N, r_a, x) = \Delta E_{\text{gpo}}(N, r_a, x) \times \{2 \times x + 6 \times (1 - x)\}, \tag{14p}$$

where  $a_1 = 3.15 \times 10^{-3}(\text{eV})$  ,  $a_2 = 5.41 \times 10^{-4}(\text{eV})$  ,  $a_3 = 2.32 \times 10^{-3}(\text{eV})$  ,  $a_4 = 4.12 \times 10^{-3}(\text{eV})$  and  $a_5 = 9.8 \times 10^{-5}(\text{eV})$ .

One also remarks that, as  $N^* = 0$ , according to the MIT,  $\Delta E_{\text{gn(gp)}}(N, r_{d(a)}, x) = 0$ .

### OPTICAL BAND GAP

Here, the optical band gap is found to be defined by:

$$E_{\text{gn1(gp1)}}(N, r_{d(a)}, x, T) \equiv E_{\text{gni(gp1)}}(r_{d(a)}, x, T) - \Delta E_{\text{gn(gp)}}(N, r_{d(a)}, x) + (-)E_{\text{Fn(Fp)}}(N, r_{d(a)}, x, T) , \tag{15}$$

where  $E_{\text{gin(gp1)}} \cdot [+E_{\text{Fn}}, -E_{\text{Fp}}] \geq 0$ , and  $\Delta E_{\text{gn(gp)}}$  are respectively determined in Equations [10, 12, 14n(p)], respectively. So, as noted above, at the MIT, Eq. (15) thus becomes:

$$E_{\text{gn1(gp1)}}(r_{d(a)}, x) = E_{\text{gno(gp0)}}(r_{d(a)}, x), \quad \text{according to: } N = N_{\text{CDn(NDp)}}(r_{d(a)}, x).$$

## OPTICAL COEFFICIENTS

The optical properties of any medium can be described by the complex refraction index  $\mathbb{N}$  and the complex dielectric function  $\varepsilon$ ,  $\mathbb{N} \equiv n - i\kappa$  and  $\varepsilon \equiv \varepsilon_1 - i\varepsilon_2$ , where  $i^2 = -1$  and  $\varepsilon \equiv \mathbb{N}^2$ . Therefore, the real and imaginary parts of  $\varepsilon$  denoted by  $\varepsilon_1$  and  $\varepsilon_2$  can thus be expressed in terms of the refraction index  $n$  and the extinction coefficient  $\kappa$  as:  $\varepsilon_1 \equiv n^2 - \kappa^2$  and  $\varepsilon_2 \equiv 2n\kappa$ . One notes that the optical absorption coefficient  $\alpha$  is related to  $\varepsilon_2$ ,  $n$ ,  $\kappa$ , and the optical conductivity  $\sigma_0$ , by<sup>[2]</sup>

$$\alpha(E, N, r_{d(a)}, \mathbf{x}, T) \equiv \frac{\hbar q^2 \times |v(E)|^2}{n(E) \times \varepsilon_{\text{free space}} \times c E} \times J(E^*) = \frac{E \times \varepsilon_2(E)}{\hbar c n(E)} \equiv \frac{2E \times \kappa(E)}{\hbar c} \equiv \frac{4\pi \sigma_0(E)}{c n(E) \times \varepsilon_{\text{free space}}},$$

$$\varepsilon_1 \equiv n^2 - \kappa^2 \text{ and } \varepsilon_2 \equiv 2n\kappa, \quad (16)$$

where, since  $E \equiv \hbar\omega$  is the photon energy, the effective photon energy:  $E^* = E - E_{\text{gn1(gp1)}}(N, r_{d(a)}, \mathbf{x}, T)$  is thus defined as the reduced photon energy.

Here,  $-q$ ,  $\hbar$ ,  $|v(E)|$ ,  $\omega$ ,  $\varepsilon_{\text{free space}}$ ,  $c$  and  $J(E^*)$  respectively represent: the electron charge, Dirac's constant, matrix elements of the velocity operator between valence (conduction)-and-conduction (valence) bands in n(p)-type semiconductors, photon frequency, permittivity of free space, velocity of light, and joint density of states. It should be noted that, if the three functions such as:  $|v(E)|^2$ ,  $J(E^*)$  and  $n(E)$  are known, then the other optical dispersion functions as those given in Eq. (16) can thus be determined. Moreover, the normal-incidence reflectance,  $R(E)$ , can be expressed in terms of  $\kappa(E)$  and  $n(E)$  as:

$$R(E, N, r_{d(a)}, \mathbf{x}, T) = \frac{[n(E)-1]^2 + \kappa(E)^2}{[n(E)+1]^2 + \kappa(E)^2}. \quad (17)$$

From Equations (16, 17), if the two optical functions,  $\varepsilon_1$  and  $\varepsilon_2$ , (or  $n$  and  $\kappa$ ), are both known, the other ones defined above can thus be determined, noting also that:  $E_{\text{gn1(gp1)}}(N, r_{d(a)}, \mathbf{x}, T) = E_{\text{gn1(gp1)}}$ , for a presentation simplicity.

Then, one has:

-at low values of  $E \gtrsim E_{\text{gn1(gp1)}}$ ,

$$J_{n(p)}(E, N, r_{d(a)}, \mathbf{x}, T) = \frac{1}{2\pi^2} \times \left(\frac{2m_r}{\hbar^2}\right)^{3/2} \times \frac{(E - E_{\text{gn1(gp1)})}^{a-(1/2)}}{E_{\text{gn1(gp1)}}^{a-1}} = \frac{1}{2\pi^2} \times \left(\frac{2m_r}{\hbar^2}\right)^{3/2} \times (E - E_{\text{gn1(gp1)})}^{1/2},$$

, for  $a=1$ , (18)

and at large values of  $E > E_{\text{gn1(gp1)}}$ ,



$$J_{n(p)}(E, N, r_{d(a)}, x, T) = \frac{1}{2\pi^2} \times \left(\frac{2m_r}{\hbar^2}\right)^{3/2} \times \frac{(E - E_{gn1(gp1)})^{a-(1/2)}}{E_{gn1(gp1)}^{a-1}} = \frac{1}{2\pi^2} \times \left(\frac{2m_r}{\hbar^2}\right)^{3/2} \times \frac{(E - E_{gn1(gp1)})^2}{E_{gn1(gp1)}^{3/2}}, \text{ for } a=5/2. \quad (19)$$

Further, one notes that, as  $E \rightarrow \infty$ , Forouhi and Bloomer (FB)<sup>[4]</sup> claimed that  $\kappa(E \rightarrow \infty) \rightarrow$  a constant, while the  $\kappa(E)$  -expressions, proposed by Van Cong<sup>[2]</sup> quickly go to 0 as  $E^{-3}$ , and consequently, their numerical results of the optical functions such as:  $\sigma_0(E)$  and  $\alpha(E)$ , given in Eq. (16), both go to 0 as  $E^{-2}$ .

Now, an improved Forouhi-Bloomer parameterization model (FB-PM), used to determine the expressions of the optical coefficients in the degenerate  $n^+(p^+) - p(n) X(x)$ - crystalline alloy, is now proposed as follows. Then, if denoting the functions  $G(E)$  and  $F(E)$  and by:

$$G(E) \equiv \sum_{i=1}^4 \frac{A_i}{E^2 - B_i E + C_i} \text{ and } F(E) \equiv \sum_{i=1}^4 \frac{A_i}{E^2 \times (1 + 10^{-4} \times \frac{E}{eV}) - B_i E + C_i}, \text{ we propose:}$$

$$\begin{aligned} \kappa(E, N, r_{d(a)}, x, T) &= G(E) \times E_{gn1(gp1)}^{3/2} \times (E^* \equiv E - E_{gn1(gp1)})^{1/2}, \text{ for } \\ E_{gn1(gp1)} &\leq E \leq 2.3 \text{ eV,} \\ &= F(E) \times (E^* \equiv E - E_{gn1(gp1)})^2, \text{ for } E \geq 2.3 \text{ eV,} \end{aligned} \quad (20)$$

being equal to 0 for  $E^* = 0$  (or for  $E = E_{gn1(gp1)}$ ), and also going to 0 as  $E^{-1}$  as  $E \rightarrow \infty$ , and further,

$$n(E, N, r_{d(a)}, x, T) = n_\infty(r_{d(a)}, x) + \sum_{i=1}^4 \frac{X_i(E_{gn1(gp1)}) \times E + Y_i(E_{gn1(gp1)})}{E^2 - B_i E + C_i}. \quad (21)$$

going to a constant as  $E \rightarrow \infty$ , since  $n(E \rightarrow \infty, r_{d(a)}, x) \rightarrow n_\infty(r_{d(a)}, x) = \sqrt{\varepsilon(r_{d(a)}, x)} \times \frac{\omega_T}{\omega_L}$ ,

$$\omega_T = 5.1 \times 10^{13} \text{ s}^{-1} \text{ [5]} \text{ and } \omega_L = 8.9755 \times 10^{13} \text{ s}^{-1}.$$

Here, the other parameters are determined by:

$$X_i(E_{gn1(gp1)}) = \frac{A_i}{Q_i} \times \left[ -\frac{B_i^2}{2} + E_{gn1(gp1)} B_i - E_{gn1(gp1)}^2 + C_i \right],$$

$$Y_i(E_{gn1(gp1)}) = \frac{A_i}{Q_i} \times \left[ \frac{B_i \times (E_{gn1(gp1)}^2 + C_i)}{2} - 2E_{gn1(gp1)} C_i \right], \quad Q_i = \frac{\sqrt{4C_i - B_i^2}}{2}, \text{ where, for } i=(1, 2, 3,$$

and 4),  $A_i = 1.154 \times A_{i(FB)} = 4.7314 \times 10^{-4}, 0.2314, 0.1118$  and  $0.0116$ ,

$B_i \equiv B_{i(FB)} = 5.871, 6.154, 9.679$  and  $13.232$ , and  $C_i \equiv C_{i(FB)} = 8.619, 9.784, 23.803$ , and  $44.119$ .

Then, as noted above, if the two optical functions,  $n$  and  $\kappa$ , are both known, the other ones defined in Equations (16, 17) can also be determined.

## NUMERICAL RESULTS

Now, some numerical results of those optical functions are investigated in the  $n(p)$ -type  $\mathbf{X(x)} \equiv \mathbf{GaTe_{1-x}P_x}$ - crystalline alloy, as follows.

### A. Metal-insulator transition (MIT)-case

As discussed above, the physical conditions used for the MIT are found to be given by:

$$T=0K, \quad N^* = 0 \quad \text{or} \quad N = N_{\text{CDn(CDp)}}, \quad \text{giving rise to:}$$

$$E_{\text{gn1(gp1)}}(N^* = 0, r_{\text{d(a)}}, x, T = 0) = E_{\text{gn1(gp1)}}(r_{\text{d(a)}}, x) = E_{\text{gno(gp0)}}(r_{\text{d(a)}}, x).$$

Then, in this MIT-case, if  $E = E_{\text{gn1(gp1)}}(r_{\text{d(a)}}, x) = E_{\text{gno(gp0)}}(r_{\text{d(a)}}, x)$ , which can be defined as the critical photon energy:  $E \equiv E_{\text{CPE}}(r_{\text{d(a)}}, x)$ , one obtains:  $\kappa_{\text{MIT}}(r_{\text{d(a)}}, x) = 0$  from Eq. (20), and from Eq. (16):  $\varepsilon_{2(\text{MIT})}(r_{\text{d(a)}}, x) = 0$ ,  $\sigma_{\text{O}(\text{MIT})}(r_{\text{d(a)}}, x) = 0$  and  $\alpha_{\text{MIT}}(r_{\text{d(a)}}, x) = 0$ , and the other functions such as:  $n_{\text{MIT}}(r_{\text{d(a)}}, x)$  from Eq. (21), and  $\varepsilon_{1(\text{MIT})}(r_{\text{d(a)}}, x)$  and  $R_{\text{MIT}}(r_{\text{d(a)}}, x)$  from Eq. (16) decrease with increasing  $r_{\text{d(a)}}$  and  $E_{\text{CPE}}$ , as those investigated in Table 1 in Appendix 1.

### B. Optical coefficients, obtained as $E \rightarrow \infty$

In Eq. (21), at any  $T$ , the choice of the real refraction index:  $n(E \rightarrow \infty, r_{\text{d(a)}}, x, T) = n_{\infty}(r_{\text{d(a)}}, x) = \sqrt{\varepsilon(r_{\text{d(a)}}, x)} \times \frac{\omega_T}{\omega_L}$ ,  $\omega_T = 5.1 \times 10^{13} \text{ s}^{-1}$  [5] and  $\omega_L = 8.9755 \times 10^{13} \text{ s}^{-1}$ , was obtained from the Lyddane-Sachs-Teller relation<sup>[5]</sup>, from which  $T(L)$  represent the transverse (longitudinal) optical phonon modes. Then, from Equations (16, 17, 20), from such the asymptotic behavior ( $E \rightarrow \infty$ ), we obtain:  $\kappa_{\infty}(r_{\text{d(a)}}, x) \rightarrow 0$  and  $\varepsilon_{2,\infty}(r_{\text{d(a)}}, x) \rightarrow 0$ , as  $E^{-1}$ , so that  $\varepsilon_{1,\infty}(r_{\text{d(a)}}, x)$ ,  $\sigma_{\text{O},\infty}(r_{\text{d(a)}}, x)$ ,  $\alpha_{\infty}(r_{\text{d(a)}}, x)$  and  $R_{\infty}(r_{\text{d(a)}}, x)$  go to their appropriate limiting constants for  $T=0K$ , as those investigated in Table 2 in Appendix 1.

### C. Variations of some optical coefficients, obtained in P(B)-X(x)-system, as functions of E

In the P(B)-X(x)-system, at  $T=0K$  and  $N = N_{\text{CDn(CDp)}}(r_{\text{P(B)}}, x)$ , our numerical results of  $n$ ,  $\kappa$ ,  $\varepsilon_1$  and  $\varepsilon_2$  are obtained from Equations (21, 20, 16), respectively, and expressed as functions of  $E [\geq E_{\text{CPE}}(r_{\text{P(B)}}, x)]$  and for given  $x$ , as those reported in Tables 3n and 3p in Appendix 1.

#### D. Variations of various optical coefficients, as functions of N

In the X(x)-system, at  $E=3.2$  eV and  $T=20$  K, for given  $r_{d(a)}$  and  $x$ , and from Equations (12, 15, 21, 20, 16), respectively, we can determine the variations of  $\eta_{n(p)} (>> 1, \text{degenerate case})$ ,  $E_{gn1(gp1)}$ ,  $n$ ,  $\kappa$ ,  $\varepsilon_1$  and  $\varepsilon_2$ , obtained as functions of N, being represented by the arrows: ↗ and ↘, as those tabulated in Tables 4n and 4p in Appendix 1.

#### E. Variations of various optical coefficients as functions of T

In the X(x)-system, at  $E=3.2$  eV and  $N = 10^{20} \text{cm}^{-3}$ , for given  $r_{d(a)}$  and  $x$ , and from Equations (12, 15, 21, 20, 16), respectively, we can determine the variations of  $\eta_{n(p)} (>> 1, \text{degenerate case})$ ,  $E_{gn1(gp1)}$ ,  $n$ ,  $\kappa$ ,  $\varepsilon_1$  and  $\varepsilon_2$ , obtained as functions of T, being represented by the arrows: ↗ and ↘, as those tabulated in Tables 5n and 5p in Appendix 1.

### CONCLUDING REMARKS

In the n(p)-type  $\mathbf{X(x)} \equiv \mathbf{GaTe_{1-x}P_x}$ - crystalline alloy, by basing on our two recent works<sup>[1,2]</sup>, for a given  $x$ , and with an increasing  $r_{d(a)}$ , the optical coefficients have been determined, as functions of the photon energy E, total impurity density N, the donor (acceptor) radius  $r_{d(a)}$ , concentration  $x$ , and temperature T.

Those results have been affected by (i) the important new  $\varepsilon(r_{d(a)}, x)$ -law, developed in Equations (8a, 8b), stating that, for a given  $x$ , due to the impurity-size effect,  $\varepsilon$  decreases (↘) with an increasing (↗)  $r_{d(a)}$ , and then by (ii) the generalized Mott critical d(a)-density defined in the metal-insulator transition (MIT),  $N_{CDn(NDp)}(r_{d(a)}, x)$ , as observed in Equations (8c, 9a).

Further, we also showed that  $N_{CDn(NDp)}$  is just the density of carriers localized in exponential band tails, with a precision of the order of  $2.91 \times 10^{-7}$ , as that given in Table 4 of Ref.<sup>[1]</sup>, according to a definition of the effective density of electrons (holes) given in parabolic conduction (valence) bands by:  $N^*(N, r_{d(a)}, x) \equiv N - N_{CDn(NDp)}(r_{d(a)}, x)$ , as defined in Eq. (9d).

In summary, due to the new  $\varepsilon(r_{d(a)}, x)$ -law and to the effective density of electrons (holes) given in parabolic conduction (valence) bands  $N^*(N, r_{d(a)}, x)$ , for a given  $x$ , and with an increasing  $r_{d(a)}$ , the numerical results of all the optical coefficients, obtained in appropriated physical conditions (E, N, T), and calculated by using Equations (15, 16, 20, 21), are reported in Tables 1, 2, 3n, 3p, 4n, 4p, 5n, and 5p in Appendix 1.

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**APPENDIX 1**

**Table 1:** In the MIT-case,  $T=0K$ ,  $N=N_{CDn(p)}(r_{d(a)}, x)$ , and the critical photon energy  $E_{CPE} = E = E_{gno(gp0)}(r_{d(a)}, x)$ , if  $E = E_{gn1(gp1)}(r_{d(a)}, x) = E_{CPE}(r_{d(a)}, x)$ , the numerical results of optical functions such as :  $n_{MIT}(r_{d(a)}, x)$ , obtained from Eq. (21), and those of other ones:  $\epsilon_{1(MIT)}(r_{d(a)}, x)$  and  $R_{MIT}(r_{d(a)}, x)$ , from Eq. (16), decrease ( $\searrow$ ) with increasing ( $\nearrow$ )  $r_{d(a)}$  and  $E_{CPE}$ .

Donor		P	Te	Sb	Sn
$r_d$ (nm) [4]	$\nearrow$	0.110	0.132	0.136	0.140
-----					
At $x=0$ ,					
$E_{CPE}$ in meV	$\nearrow$	1791.7	1796	1796.1	1796.6
$n_{MIT}$	$\searrow$	3.315	3.178	3.174	3.161
$\epsilon_{1(MIT)}$	$\searrow$	10.99	10.10	10.07	9.99
$R_{MIT}$	$\searrow$	0.288	0.272	0.271	0.2697
-----					
At $x=0.5$ ,					
$E_{CPE}$ in meV	$\nearrow$	1792.1	1796	1796.1	1796.6
$n_{MIT}$	$\searrow$	3.263	3.129	3.125	3.112
$\epsilon_{1(MIT)}$	$\searrow$	10.64	9.79	9.76	9.68
$R_{MIT}$	$\searrow$	0.282	0.266	0.265	0.264
-----					
At $x=1$ ,					
$E_{CPE}$ in meV	$\nearrow$	1792.7	1796	1796.1	1796.5
$n_{MIT}$	$\searrow$	3.208	3.078	3.074	3.062
$\epsilon_{1(MIT)}$	$\searrow$	10.29	9.47	9.45	9.38
$R_{MIT}$	$\searrow$	0.275	0.2597	0.2592	0.2577
-----					
Acceptor		B	Ga	In	Cd
$r_a$ (nm)	$\nearrow$	0.088	0.126	0.144	0.148
-----					
At $x=0$ ,					
$E_{CPE}$ in meV	$\nearrow$	1770.5	1796	1803	1807
$n_{MIT}$	$\searrow$	3.917	3.178	3.086	3.045

$\varepsilon_{1(MIT)}$	↘	15.34	10.10	9.522	9.271
$R_{MIT}$	↘	0.352	0.272	0.261	0.255

At  $x=0.5$ ,

$E_{CPE}$ in meV	↗	1764.2	1796	1805	1809.4
$n_{MIT}$	↘	3.853	3.129	3.038	2.997
$\varepsilon_{1(MIT)}$	↘	14.85	9.79	9.23	8.98
$R_{MIT}$	↘	0.346	0.266	0.255	0.250

At  $x=1$ ,

$E_{CPE}$ in meV	↗	1756.8	1796	1807	1812.5
$n_{MIT}$	↘	3.789	3.078	2.988	2.948
$\varepsilon_{1(MIT)}$	↘	14.36	9.47	8.93	8.69
$R_{MIT}$	↘	0.339	0.260	0.248	0.243

**Table 2:** Here, as  $T=0K$  and  $N=N_{CDn(p)}(r_{d(a)}, x)$ , and for  $E \rightarrow \infty$  the numerical results of  $n_{\infty}(r_{d(a)}, x)$ ,  $\varepsilon_{1,\infty}(r_{d(a)}, x)$ ,  $\sigma_{0,\infty}(r_{d(a)}, x)$ ,  $\alpha_{\infty}(r_{d(a)}, x)$  and  $R_{\infty}(r_{d(a)}, x)$  go to their appropriate limiting constants.

Donor		P	Te	Sb	Sn
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At  $x=0$ ,

$n_{\infty}$	↘	2.1277	1.9928	1.9886	1.9762
$\varepsilon_{1,\infty}$	↘	4.5269	3.9712	3.9547	3.9053
$\sigma_{0,\infty}$ in $\frac{10^5}{\Omega \times cm}$	↘	9.7087	9.0933	9.0743	9.0174
$\alpha_{\infty}$ in $(10^9 \times cm^{-1})=$		2.1602			
$R_{\infty}$	↘	0.1300	0.1100	0.1094	0.1076

At  $x=0.5$ ,

$n_{\infty}$	↘	2.0751	1.9436	1.9395	1.9274
$\varepsilon_{1,\infty}$	↘	4.3061	3.7775	3.7618	3.7148



$\sigma_{0,\infty}$ in $\frac{10^5}{\Omega \times cm}$	↘	9.4689	8.8687	8.8502	8.7947
$\alpha_\infty$ in $(10^9 \times cm^{-1})$		2.1602			
$R_\infty$	↘	0.1222	0.1027	0.1021	0.1003

At  $x=1$ ,

$n_\infty$	↘	2.0212	1.8931	1.8891	1.8773
$\varepsilon_{1,\infty}$	↘	4.0853	3.5838	3.5689	3.5343
$\sigma_{0,\infty}$ in $\frac{10^5}{\Omega \times cm}$	↘	9.2229	8.6383	8.6203	8.5663
$\alpha_\infty$ in $(10^9 \times cm^{-1})$		2.1602			
$R_\infty$	↘	0.1142	0.0953	0.0947	0.0930

Acceptor	B	Ga	In	Cd
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At  $x=0$ ,

$n_\infty$	↘	2.716	1.993	1.905	1.866
$\varepsilon_{1,\infty}$	↘	7.374	3.971	3.629	3.483
$\sigma_{0,\infty}$ in $\frac{10^5}{\Omega \times cm}$	↘	12.39	9.093	8.693	8.517
$\alpha_\infty$ in $(10^9 \times cm^{-1})$		2.160			
$R_\infty$	↘	0.213	0.110	0.097	0.091

At  $x=0.5$ ,

$n_\infty$	↘	2.648	1.943	1.858	1.820
$\varepsilon_{1,\infty}$	↘	7.014	3.777	3.452	3.314
$\sigma_{0,\infty}$ in $\frac{10^5}{\Omega \times cm}$	↘	12.08	8.869	8.478	8.306
$\alpha_\infty$ in $(10^9 \times cm^{-1})$		2.160			
$R_\infty$	↘	0.204	0.103	0.090	0.085

At  $x=1$ ,

$n_\infty$	↘	2.580	1.893	1.810	1.773
$\varepsilon_{1,\infty}$	↘	6.655	3.584	3.275	3.144
$\sigma_{0,\infty}$ in $\frac{10^5}{\Omega \times cm}$	↘	11.77	8.64	8.26	8.09

$$\alpha_{\infty} \text{ in } (10^9 \times \text{cm}^{-1})=2.160$$

$$R_{\infty} \quad \searrow \quad 0.195 \quad 0.095 \quad 0.083 \quad 0.078$$

**Table 3n:** In the P-X(x)-system, and at T=0K and  $N = N_{CDn}(r_p, x)$ , according to the MIT, our numerical results of  $n$ ,  $\kappa$ ,  $\epsilon_1$  and  $\epsilon_2$  are obtained from Equations (21, 20, 16), respectively, and expressed as functions of  $E [\geq E_{CPE}(r_p, x)]$  and  $x$ , noting that (i)  $\kappa = 0$  and  $\epsilon_2 = 0$  at  $E = E_{CPE}(r_p, x)$ , and  $\kappa \rightarrow 0$  and  $\epsilon_2 \rightarrow 0$  as  $E \rightarrow \infty$ .

E in eV	$n$	$\kappa$	$\epsilon_1$	$\epsilon_2$
<b>At x=0,</b>				
<b><math>E_{CPE} = 1.7917</math></b>	<b>3.3156</b>	<b>0</b>	<b>10.993</b>	<b>0</b>
2	3.462	0.187	11.949	1.296
2.5	3.992	0.190	15.898	1.516
3	4.176	1.199	15.999	10.016
3.5	3.640	1.520	10.938	11.062
4	3.771	1.476	12.044	11.134
4.5	4.085	2.387	10.992	19.502
5	2.609	3.441	-5.032	17.955
5.5	1.534	2.487	-3.832	7.632
6	1.616	1.888	-0.952	6.103
...				
<b><math>10^{22}</math></b>	<b>2.1277</b>	<b>0</b>	<b>4.5269</b>	<b>0</b>

<b>At x=0.5,</b>				
<b><math>E_{CPE} = 1.7921</math></b>	<b>3.7632</b>	<b>0</b>	<b>10.6455</b>	<b>0</b>
2	3.408	0.187	11.583	1.276
2.5	3.938	0.190	15.474	1.494
3	4.123	1.198	15.559	9.881
3.5	3.587	1.519	10.559	10.896
4	3.718	1.475	11.650	10.974
4.5	4.032	2.386	10.567	19.244
5	2.557	3.440	-5.296	17.590

5.5	1.482	2.487	-3.987	7.371
6	1.564	1.887	-1.116	5.904
...				
<b>10<sup>22</sup></b>	<b>2.0751</b>	<b>0</b>	<b>4.3061</b>	<b>0</b>

At x=1,

<b>E<sub>CPE</sub> =1.7927</b>	<b>3.2084</b>	<b>0</b>	<b>10.2943</b>	<b>0</b>
2	3.354	0.187	11.213	1.254
2.5	3.883	0.189	15.044	1.471
3	4.068	1.197	15.114	9.741
3.5	3.533	1.518	10.177	10.724
4	3.664	1.475	11.253	10.808
4.5	3.978	2.385	10.137	18.977
5	2.503	3.439	-5.558	17.214
5.5	1.429	2.486	-4.138	7.103
6	1.511	1.887	-1.279	5.701
...				
<b>10<sup>22</sup></b>	<b>2.0212</b>	<b>0</b>	<b>4.0853</b>	<b>0</b>

E in eV	<i>n</i>	<i>κ</i>	$\epsilon_1$	$\epsilon_2$
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**Table 3p:** In the B-X(x)-system, and at T=0K and  $N = N_{CDP}(r_B, x)$ , according to the MIT, our numerical results of *n*, *κ*,  $\epsilon_1$  and  $\epsilon_2$  are obtained from Equations (21, 20, 16), respectively, and expressed as functions of  $E [\geq E_{CPE}(r_B, x)]$  and x, noting that (i)  $\kappa = 0$  and  $\epsilon_2 = 0$  at  $E = E_{CPE}(r_B, x)$ , and  $\kappa \rightarrow 0$  and  $\epsilon_2 \rightarrow 0$  as  $E \rightarrow \infty$

E in eV	<i>n</i>	<i>κ</i>	$\epsilon_1$	$\epsilon_2$
At x=0,				
<b>E<sub>CPE</sub> =1.7705</b>	<b>3.9167</b>	<b>0</b>	<b>15.3404</b>	<b>0</b>
2	4.079	0.193	16.606	1.575
2.5	4.620	0.201	21.306	1.862
3	4.794	1.242	21.439	11.906
3.5	4.235	1.557	15.513	13.194

4	4.368	1.505	16.814	13.144
4.5	4.685	2.424	16.073	22.717
5	3.187	3.486	-1.995	22.226
5.5	2.102	2.516	-1.912	10.575
6	2.187	1.907	1.146	8.341
...				
<b>10<sup>22</sup></b>	<b>2.7156</b>	<b>0</b>	<b>7.3743</b>	<b>0</b>

At  $x=0.5$ ,

<b><math>E_{CPE} = 1.7642</math></b>	<b>3.8535</b>	<b>0</b>	<b>14.8494</b>	<b>0</b>
2	4.021	0.195	16.133	1.565
2.5	4.565	0.205	20.799	1.871
3	4.736	1.254	20.853	11.881
3.5	4.171	1.569	14.934	13.086
4	4.303	1.530	16.229	13.022
4.5	4.621	2.435	15.427	22.511
5	3.118	3.500	-2.529	21.823
5.5	2.028	2.524	-2.256	10.241
6	2.115	1.912	0.814	8.089
...				
<b>10<sup>22</sup></b>	<b>2.6485</b>	<b>0</b>	<b>7.0146</b>	<b>0</b>

At  $x=1$ ,

<b><math>E_{CPE} = 1.7568</math></b>	<b>3.7893</b>	<b>0</b>	<b>14.3590</b>	<b>0</b>
2	3.963	0.196	15.668	1.557
2.5	4.511	0.209	20.304	1.886
3	4.677	1.269	20.268	11.876
3.5	4.105	1.582	14.344	12.989
4	4.237	1.523	15.636	12.908
4.5	4.557	2.449	14.769	22.317
5	3.045	3.516	-3.087	21.416
5.5	1.952	2.534	-2.610	9.896
6	2.040	1.919	0.477	7.830

...

**10<sup>22</sup>**                      **2.5797**                      **0**                      **6.6548**                      **0**

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E in eV	<i>n</i>	<i>κ</i>	<i>ε</i> <sub>1</sub>	<i>ε</i> <sub>2</sub>
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**Table 4n:** In the X(x)-system, at E=3.2 eV and T=20 K, for given r<sub>d</sub> and x, and from Equations (12, 15, 21, 20, 16), respectively, we can determine the variations of η<sub>n</sub>( $\gg 1$ , degenerate case), E<sub>gn1</sub>, n, κ, ε<sub>1</sub> and ε<sub>2</sub>, obtained as functions of N, being represented by the arrows: ↗ and ↘, noting that both η<sub>n</sub> and E<sub>gn1</sub> increase with increasing N.

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N (10 <sup>18</sup> cm <sup>-3</sup> )	↗ 15	26	60	100
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**x=0**

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For r<sub>d</sub> = r<sub>Te</sub>,

η <sub>n</sub> $\gg 1$	↗	234	340	599	844
E <sub>gn1</sub> in eV	↗	1.672	1.711	1.854	2.026

---

n	↘	3.994	3.956	3.815	3.641
κ	↘	1.730	1.643	1.342	1.022
ε <sub>1</sub>	↘	12.958	12.951	12.751	12.210
ε <sub>2</sub>	↘	13.820	13.000	10.241	7.445

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For r<sub>d</sub> = r<sub>Sb</sub>,

η <sub>n</sub> $\gg 1$	↗	233	340	599	844
E <sub>gn1</sub> in eV	↗	1.673	1.713	1.857	2.029

---

n	↘	3.989	3.950	3.808	3.633
κ	↘	1.728	1.640	1.337	1.016
ε <sub>1</sub>	↘	12.925	12.918	12.713	12.166
ε <sub>2</sub>	↘	13.783	12.954	10.186	7.387

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For r<sub>d</sub> = r<sub>Sn</sub>,

η <sub>n</sub> $\gg 1$	↗	233	340	599	844
E <sub>gn1</sub> in eV	↗	1.676	1.717	1.864	2.039

---

n	↘	3.973	3.933	3.788	3.610
$\kappa$	↘	1.720	1.630	1.322	0.999
$\varepsilon_1$	↘	12.826	12.817	12.601	12.036
$\varepsilon_2$	↘	13.671	12.821	10.020	7.216

**x=0.5**

For  $r_d = r_{Te}$ ,

$\eta_n \gg 1$	↗	129	187	328	462
$E_{gn1}$ in eV	↘	1.597	1.579	1.575	↗ 1.599
n	↗	4.017	4.034	4.038	↘ 4.015
$\kappa$	↗	1.906	1.948	1.958	↘ 1.901
$\varepsilon_1$	↘	12.508	12.480	12.473	↗ 12.510
$\varepsilon_2$	↗	15.311	15.715	15.818	↘ 15.264

For  $r_d = r_{Sb}$ ,

$\eta_n \gg 1$	↗	129	187	328	462
$E_{gn1}$ in eV	↘	1.598	1.580	1.577	↗ 1.602
n	↗	4.012	4.029	4.032	↘ 4.008
$\kappa$	↗	1.903	1.944	1.953	↘ 1.894
$\varepsilon_1$	↘	12.476	12.450	12.444	↗ 12.482
$\varepsilon_2$	↗	15.273	15.668	15.751	↘ 15.182

For  $r_d = r_{Sn}$ ,

$\eta_n \gg 1$	↗	129	187	328	462
$E_{gn1}$ in eV	↘	1.601	1.584	1.583	↗ 1.610
n	↗	3.997	4.013	4.014	↘ 3.988
$\kappa$	↗	1.896	1.935	1.937	↘ 1.873
$\varepsilon_1$	↘	12.383	12.359	12.357	↗ 12.396
$\varepsilon_2$	↗	15.161	15.527	15.552	↘ 14.940

**x=1**



For  $r_d = r_{Te}$ ,

$\eta_n \gg 1$	$\nearrow$	93	135	236	331
$E_{gn1}$ in eV	$\searrow$	1.678	1.671 $\nearrow$	1.678	1.704
n	$\nearrow$	3.889	3.896 $\searrow$	3.888	3.863
$\kappa$	$\nearrow$	1.718	1.734 $\searrow$	1.716	1.658
$\varepsilon_1$	$\searrow$	12.173	12.171 $\nearrow$	12.173	12.174
$\varepsilon_2$	$\nearrow$	13.364	13.511 $\searrow$	13.346	12.808

For  $r_d = r_{Sb}$ ,

$\eta_n \gg 1$	$\nearrow$	93	135	236	331
$E_{gn1}$ in eV	$\searrow$	1.678	1.671 $\nearrow$	1.680	1.707
n	$\nearrow$	3.884	3.891 $\searrow$	3.883	3.857
$\kappa$	$\nearrow$	1.716	1.732 $\searrow$	1.713	1.653
$\varepsilon_1$	$\searrow$	12.143	12.141 $\nearrow$	12.143	12.1431
$\varepsilon_2$	$\nearrow$	13.336	13.477 $\searrow$	13.300	12.754

For  $r_d = r_{Sn}$ ,

$\eta_n \gg 1$	$\nearrow$	93	135	236	331
$E_{gn1}$ in eV		1.680 $\searrow$	1.674 $\nearrow$	1.684	1.713
n		3.871 $\nearrow$	3.876 $\searrow$	3.866	3.839
$\kappa$		1.712 $\nearrow$	1.725 $\searrow$	1.702	1.640
$\varepsilon_1$		12.051 $\searrow$	12.050 $\nearrow$	12.052 $\searrow$	12.051
$\varepsilon_2$	$\nearrow$	13.254	13.377 $\searrow$	13.165	12.593

N ( $10^{18} \text{ cm}^{-3}$ )	$\nearrow$	15	26	60	100
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**Table 4p:** In the X(x)-system, at E=3.2 eV and T=20 K, for given  $r_d$  and x, and from Equations (12, 15, 21, 20, 16), respectively, we can determine the variations of  $\eta_p (\gg 1, \text{degenerate case})$ ,  $E_{gp1}$ , n,  $\kappa$ ,  $\varepsilon_1$  and  $\varepsilon_2$ , obtained as functions of N, being represented by the arrows: ↗ and ↘, noting that both  $\eta_p$  and  $E_{gp1}$  increase with increasing N.

N ( $10^{18} \text{ cm}^{-3}$ )	↗	15	26	60	100
<b>x=0</b>					
-----					
For $r_a = r_{Ga}$ ,					
$\eta_p \gg 1$	↗	199	312	578	826
$E_{gp1}$ in eV	↗	2.001	2.150	2.516	2.871
n	↘	3.666	3.511	3.112	2.723
$\kappa$	↘	1.066	0.818	0.346	0.080
$\varepsilon_1$	↘	12.306	11.661	9.566	7.302
$\varepsilon_2$	↘	7.820	5.743	2.155	0.434
-----					
For $r_a = r_{In}$ ,					
$\eta_p \gg 1$	↗	186	302	570	820
$E_{gp1}$ in eV	↗	1.997	2.152	2.528	2.889
n	↘	3.582	3.421	3.012	2.594
$\kappa$	↘	1.073	0.814	0.335	0.072
$\varepsilon_1$	↘	11.683	11.042	8.958	6.725
$\varepsilon_2$	↘	7.688	5.573	2.016	0.372
-----					
For $r_a = r_{Cd}$ ,					
$\eta_p \gg 1$	↗	178	296	566	816
$E_{gp1}$ in eV	↗	1.993	2.151	2.532	2.896
n	↘	3.548	3.383	2.969	2.547
$\kappa$	↘	1.079	0.815	0.331	0.069
$\varepsilon_1$	↘	11.421	10.782	8.703	6.485

$\epsilon_2$	$\searrow$	7.659	5.517	1.965	0.349
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**x=0.5**

For  $r_a = r_{Ga}$ ,

$\eta_p \gg 1$	$\nearrow$	93	158	307	445
$E_{gp1}$ in eV	$\nearrow$	1.888	1.973	2.179	2.376

n	$\searrow$	3.732	3.645	3.431	3.219
$\kappa$	$\searrow$	1.277	1.115	0.772	0.503
$\epsilon_1$	$\searrow$	12.299	12.044	11.175	10.108
$\epsilon_2$	$\searrow$	9.530	8.133	5.301	3.240

For  $r_a = r_{In}$ ,

$\eta_p \gg 1$	$\nearrow$	80	149	300	439
$E_{gp1}$ in eV	$\nearrow$	1.882	1.973	2.185	2.386

n	$\searrow$	3.652	3.560	3.339	3.122
$\kappa$	$\searrow$	1.287	1.116	0.763	0.491
$\epsilon_1$	$\searrow$	11.680	11.428	10.565	9.505
$\epsilon_2$	$\searrow$	9.401	7.946	5.096	3.064

For  $r_a = r_{Cd}$ ,

$\eta_p \gg 1$	$\nearrow$	72	143	296	435
$E_{gp1}$ in eV	$\nearrow$	1.878	1.972	2.187	2.390

n	$\searrow$	3.619	3.524	3.299	3.080
$\kappa$	$\searrow$	1.295	1.118	0.760	0.486
$\epsilon_1$	$\searrow$	11.416	11.166	10.305	9.249
$\epsilon_2$	$\searrow$	9.377	7.883	5.015	2.993

**x=1**

For  $r_a = r_{Ga}$ ,

$\eta_p \gg 1$	$\nearrow$	47	99	210	310
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$E_{gp1}$ in eV	↗	1.854	1.930	2.099	2.255
$n$	↘	3.716	3.638	3.464	3.299
$\kappa$	↘	1.343	1.195	0.898	0.661
$\varepsilon_1$	↘	12.001	11.811	11.196	10.447
$\varepsilon_2$	↘	9.983	8.696	6.221	4.364
-----					
For $r_a = r_{In}$ ,					
$\eta_p \gg 1$	↗	28	87	202	303
$E_{gp1}$ in eV	↗	1.838	1.924	2.100	2.259
$n$	↘	3.648	3.561	3.380	3.211
$\kappa$	↘	1.376	1.207	0.897	0.656
$\varepsilon_1$	↘	11.418	11.228	10.622	9.884
$\varepsilon_2$	↘	10.039	8.597	6.063	4.213
-----					
For $r_a = r_{Cd}$ ,					
$\eta_p \gg 1$	↗	13	80	197	300
$E_{gp1}$ in eV	↗	1.824	1.919	2.100	2.260
$n$	↘	3.626	3.530	3.344	3.173
$\kappa$	↘	1.404	1.216	0.898	0.654
$\varepsilon_1$	↘	11.174	10.980	10.378	9.643
$\varepsilon_2$	↘	10.184	8.582	6.004	4.152
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$N$ ( $10^{18} \text{ cm}^{-3}$ )	↗	15	26	60	100

**Table 5n:** In the X(x)-system, at  $E=3.2 \text{ eV}$  and  $N = 10^{20} \text{ cm}^{-3}$ , for given  $r_d$  and  $x$ , and from Equations (12, 15, 21, 20, 16), respectively, we can determine the variations of  $\eta_n$  ( $\gg 1$ , degenerate case),  $E_{gn1}$ ,  $n$ ,  $\kappa$ ,  $\varepsilon_1$  and  $\varepsilon_2$ , obtained as functions of  $T$ , being represented by the arrows: ↗ and ↘, noting that both  $\eta_n$  and  $E_{gn1}$  decrease with increasing  $T$ .

$T$ in K	↗	20	50	100	300
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$x=0$

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For  $r_d = r_{Te}$ ,

$\eta_n \gg 1$	$\searrow$	844	338	169	56
$E_{gn1}$ in eV	$\searrow$	2.026	2.021	2.011	1.956
n	$\nearrow$	3.641	3.645	3.656	3.712
$\kappa$	$\nearrow$	1.022	1.030	1.048	1.147
$\varepsilon_1$	$\nearrow$	12.210	12.227	12.267	12.463
$\varepsilon_2$	$\nearrow$	7.445	7.508	7.665	8.512

For  $r_d = r_{Sb}$ ,

$\eta_n \gg 1$	$\searrow$	844	338	169	56
$E_{gn1}$ in eV	$\searrow$	2.029	2.025	2.014	1.960
n	$\nearrow$	3.633	3.638	3.648	3.704
$\kappa$	$\nearrow$	1.016	1.024	1.042	1.140
$\varepsilon_1$	$\nearrow$	12.166	12.183	12.224	12.421
$\varepsilon_2$	$\nearrow$	7.387	7.449	7.606	8.449

For  $r_d = r_{Sn}$ ,

$\eta_n \gg 1$	$\searrow$	844	337	169	56
$E_{gn1}$ in eV	$\searrow$	2.039	2.035	2.024	1.970
n	$\nearrow$	3.610	3.615	3.626	3.682
$\kappa$	$\nearrow$	0.999	1.007	1.025	1.122
$\varepsilon_1$	$\nearrow$	12.036	12.053	12.095	12.295
$\varepsilon_2$	$\nearrow$	7.216	7.278	7.432	8.263

**x=0.5**

For  $r_d = r_{Te}$ ,

$\eta_n \gg 1$	$\searrow$	462	185	92.4	31
$E_{gn1}$ in eV	$\searrow$	1.599	1.594	1.583	1.516
n	$\nearrow$	4.015	4.019	4.030	4.094
$\kappa$	$\nearrow$	1.901	1.911	1.938	2.103
$\varepsilon_1$	$\searrow$	12.510	12.504	12.487	12.338

$\varepsilon_2$	$\nearrow$	15.264	15.362	15.626	17.216
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For  $r_d = r_{Sb}$ ,

$\eta_n \gg 1$	$\searrow$	462	185	92.4	31
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$E_{gn1}$ in eV	$\searrow$	1.601	1.597	1.586	1.519
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n	$\nearrow$	4.008	4.013	4.024	4.087
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$\kappa$	$\nearrow$	1.894	1.904	1.931	2.095
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$\varepsilon_1$	$\searrow$	12.482	12.476	12.459	12.313
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$\varepsilon_2$	$\nearrow$	15.182	15.280	15.543	17.128
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For  $r_d = r_{Sn}$ ,

$\eta_n \gg 1$	$\searrow$	462	185	92.4	31
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$E_{gn1}$ in eV	$\searrow$	1.610	1.606	1.595	1.527
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n	$\nearrow$	3.988	3.992	4.003	4.067
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$\kappa$	$\nearrow$	1.873	1.883	1.911	2.074
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$\varepsilon_1$	$\searrow$	12.396	12.390	12.375	12.238
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$\varepsilon_2$	$\nearrow$	14.940	15.037	15.297	16.866
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**x=1**

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For  $r_d = r_{Te}$ ,

$\eta_n \gg 1$	$\searrow$	331	133	66	22
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$E_{gn1}$ in eV	$\searrow$	1.704	1.700	1.688	1.608
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n	$\nearrow$	3.863	3.867	3.879	3.956
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$\kappa$	$\nearrow$	1.658	1.667	1.695	1.879
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$\varepsilon_1$	$\nearrow$	12.1739	12.1745	12.1747 $\searrow$	12.1205
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$\varepsilon_2$	$\nearrow$	12.808	12.897	13.155	14.864
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For  $r_d = r_{Sb}$ ,

$\eta_n \gg 1$	$\searrow$	331	133	66	22
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$E_{gn1}$ in eV	$\searrow$	1.707	1.702	1.690	1.610
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n	$\nearrow$	3.857	3.861	3.873	3.950
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$\kappa$	$\nearrow$	1.653	1.663	1.691	1.874
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$\epsilon_1$	$\nearrow$	12.1431	12.1438	12.1442	$\searrow$	12.0919
$\epsilon_2$	$\nearrow$	12.754	12.843	13.100		14.804

For  $r_d = r_{Sn}$ ,

$\eta_n \gg 1$	$\searrow$	331	133	66		22
$E_{gn1}$ in eV	$\searrow$	1.713	1.708	1.696		1.616
n	$\nearrow$	3.839	3.843	3.856		3.932
$\kappa$	$\nearrow$	1.640	1.650	1.677		1.860
$\epsilon_1$	$\nearrow$	12.0506	12.0516	12.0528	$\searrow$	12.0061
$\epsilon_2$	$\nearrow$	12.593	12.681	12.936		14.627
T in K	$\nearrow$	20	50	100		300

**Table 5p:** In the X(x)-system, at  $E=3.2$  eV and  $N = 10^{20} \text{cm}^{-3}$ , for given  $r_a$  and x, and from Equations (12, 15, 21, 20, 16), respectively, we can determine the variations of  $\eta_p (\gg 1, \text{degenerate case})$ ,  $E_{gp1}$ , n,  $\kappa$ ,  $\epsilon_1$  and  $\epsilon_2$ , obtained as functions of T, being represented by the arrows:  $\nearrow$  and  $\searrow$ , noting that both  $\eta_p$  and  $E_{gp1}$  decrease with increasing T.

T in K	$\nearrow$	20	50	100		300
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**x=0**

For  $r_a = r_{Ga}$ ,

$\eta_p \gg 1$	$\searrow$	826	331	165		55
$E_{gp1}$ in eV	$\searrow$	2.871	2.866	2.856		2.801
n	$\nearrow$	2.703	2.708	2.721		2.785
$\kappa$	$\nearrow$	0.080	0.082	0.088		0.118
$\epsilon_1$	$\nearrow$	7.302	7.329	7.396		7.744
$\epsilon_2$	$\nearrow$	0.434	0.446	0.477		0.656

For  $r_a = r_{In}$ ,

$\eta_p \gg 1$	$\searrow$	820	328	164		55
$E_{gp1}$ in eV	$\searrow$	2.889	2.884	2.874		2.819

$n$	$\nearrow$	2.594	2.599	2.612	2.676
$\kappa$	$\nearrow$	0.072	0.074	0.079	0.107
$\varepsilon_1$	$\nearrow$	6.725	6.751	6.816	7.151
$\varepsilon_2$	$\nearrow$	0.372	0.383	0.411	0.574

For  $r_a = r_{Cd}$ ,

$\eta_p \gg 1$	$\searrow$	816	327	163	54
$E_{gp1}$ in eV	$\searrow$	2.896	2.891	2.881	2.826

$n$	$\nearrow$	2.547	2.552	2.565	2.629
$\kappa$	$\nearrow$	0.069	0.070	0.075	0.103
$\varepsilon_1$	$\nearrow$	6.485	6.510	6.573	6.904
$\varepsilon_2$	$\nearrow$	0.349	0.360	0.387	0.544

$x=0.5$

For  $r_a = r_{Ga}$ ,

$\eta_p \gg 1$	$\searrow$	445	178	89	30
$E_{gp1}$ in eV	$\searrow$	2.376	2.372	2.360	2.293

$n$	$\nearrow$	3.219	3.223	3.236	3.309
$\kappa$	$\nearrow$	0.503	0.509	0.523	0.610
$\varepsilon_1$	$\nearrow$	10.108	10.133	10.200	10.578
$\varepsilon_2$	$\nearrow$	3.240	3.279	3.384	4.036

For  $r_a = r_{In}$ ,

$\eta_p \gg 1$	$\searrow$	439	175	88	29
$E_{gp1}$ in eV	$\searrow$	2.386	2.382	2.370	2.303

$n$	$\nearrow$	3.122	3.127	3.139	3.212
$\kappa$	$\nearrow$	0.491	0.496	0.510	0.596
$\varepsilon_1$	$\nearrow$	9.505	9.530	9.595	9.964
$\varepsilon_2$	$\nearrow$	3.064	3.101	3.202	3.829

For  $r_a = r_{Cd}$ ,

$\eta_p \gg 1$	↘	435	174	87	29
$E_{gp1}$ in eV	↘	2.390	2.386	2.374	2.307
$n$	↗	3.080	3.083	3.097	3.170
$\kappa$	↗	0.486	0.491	0.505	0.591
$\varepsilon_1$	↗	9.249	9.273	9.337	9.702
$\varepsilon_2$	↗	2.993	3.030	3.129	3.745

**x=1**

For  $r_a = r_{Ga}$ ,

$\eta_p \gg 1$	↘	310	124	62	21
$E_{gp1}$ in eV	↘	2.255	2.251	2.238	2.159
$n$	↗	3.299	3.304	3.317	3.402
$\kappa$	↗	0.661	0.667	0.685	0.803
$\varepsilon_1$	↗	10.447	10.470	10.535	10.928
$\varepsilon_2$	↗	4.364	4.411	4.547	5.467

For  $r_a = r_{In}$ ,

$\eta_p \gg 1$	↘	303	121	61	20
$E_{gp1}$ in eV	↘	2.259	2.255	2.242	2.163
$n$	↗	3.211	3.216	3.230	3.314
$\kappa$	↗	0.656	0.662	0.680	0.797
$\varepsilon_1$	↗	9.884	9.906	9.969	10.350
$\varepsilon_2$	↗	4.213	4.258	4.390	5.287

For  $r_a = r_{Cd}$ ,

$\eta_p \gg 1$	↘	300	120	60	20
$E_{gp1}$ in eV	↘	2.260	2.256	2.244	2.164
$n$	↗	3.173	3.178	3.192	3.276
$\kappa$	↗	0.654	0.660	0.678	0.796
$\varepsilon_1$	↗	9.643	9.665	9.727	10.102
$\varepsilon_2$	↗	4.152	4.197	4.328	5.214

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T in K	↗	20	50	100	300
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